Lattice Effective Field Theory for Nuclei from A = 4 to A = 28

Nuclear Lattice EFT Collaboration

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HELMHOLTZ





<u>Outline</u>

Brief introduction to Lattice EFT for nuclei

Carbon-12 and the Hoyle state

Production of Carbon-12 in red giant stars

Bounds on the anthropic scenario

Preliminary results up to A = 28

Low-energy nucleons

Chiral effective field theory on the lattice ...



Current status of lattice chiral EFT Improved NNLO interaction ...

Epelbaum, Hammer, Meißner, Rev. Mod. Phys. 81 (2009) 1773



$$\begin{split} \mathcal{A}_{\mathrm{LO}} &= C_{S=0,I=1} f(\boldsymbol{q}) \left(\frac{1}{4} - \frac{1}{4} \,\boldsymbol{\sigma}_i \cdot \boldsymbol{\sigma}_j \right) \left(\frac{3}{4} + \frac{1}{4} \,\boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \right) \\ &+ C_{S=1,I=0} f(\boldsymbol{q}) \left(\frac{3}{4} + \frac{1}{4} \,\boldsymbol{\sigma}_i \cdot \boldsymbol{\sigma}_j \right) \left(\frac{1}{4} - \frac{1}{4} \,\boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \right) \\ &- \tilde{g}_{\pi N}^2 \boldsymbol{\tau}_i \cdot \boldsymbol{\tau}_j \frac{\boldsymbol{\sigma}_i \cdot \boldsymbol{q} \,\boldsymbol{\sigma}_j \cdot \boldsymbol{q}}{\boldsymbol{q}^2 + M_{\pi}^2} \,, \end{split}$$

Smearing of LO contact interactions

$$\begin{split} C_0 &= \frac{3}{4} \, C_{S=0,I=1} + \frac{1}{4} \, C_{S=1,I=0} \\ C_I &= \frac{1}{4} \, C_{S=0,I=1} - \frac{1}{4} \, C_{S=1,I=0} \end{split}$$

Fix constants from the two-nucleon sector Lattice EFT is predictive for A > 2 ...

$$\cos \delta_L \cdot j_L(kR_{wall}) = \sin \delta_L \cdot y_L(kR_{wall}),$$

$$\delta_L = \tan^{-1} \left[\frac{j_L(kR_{wall})}{y_L(kR_{wall})} \right].$$





Gaussian smearing of contact terms

Borasoy, Krebs, Lee, Meißner, Nucl. Phys. A768 (2006) 179; Eur. Phys. J. A31 (2007) 105; Lee, Prog. Part. Nucl. Phys. 63 (2009) 179 Euclidean time projection Ground state energy ...

$$Z_A(t) = \langle \psi_A | \exp(-tH) | \psi_A \rangle$$

Lattice Hamiltonian (discretized)

Choice of trial wavefunction:

- Standing waves
- Alpha clusters
- Shell model wavefunctions



Operator expectation values ...

 $E_A = -\lim_{t \to \infty}$

$$Z_A^{\mathcal{O}}(t) = \langle \psi_A | \exp(-tH/2) \mathcal{O} \exp(-tH/2) | \psi_A \rangle$$

$$\lim_{t \to \infty} \frac{Z_A^{\mathcal{O}}(t)}{Z_A(t)} = \langle \psi_A | \mathcal{O} | \psi_A \rangle$$

Decoupling of nucleon-nucleon interactions Hubbard-Stratonovich transformation ...

$$\exp\left[-\frac{C}{2}(N^{\dagger}N)^{2}\right] = \sqrt{\frac{1}{2\pi}} \int_{-\infty}^{\infty} ds \exp\left[-\frac{1}{2}s^{2} + \sqrt{-C}s(N^{\dagger}N)\right]$$



Auxiliary Field Quantum Monte Carlo (AFQMC) Discretized Euclidean time evolution ...

$$= M_{\rm LO}$$
 $= M_{\rm approx}$ $= O_{\rm observable}$

sign problem (not too severe)

positive definite strength can be varied

Hybrid Monte Carlo sampling

$$e^{-E_{0,\mathrm{LO}}a_{t}} = \lim_{n_{t}\to\infty} Z_{n_{t}+1,\mathrm{LO}}/Z_{n_{t},\mathrm{LO}}$$
$$\langle O \rangle_{0,\mathrm{LO}} = \lim_{n_{t}\to\infty} Z_{n_{t},\mathrm{LO}}^{\langle O \rangle}/Z_{n_{t},\mathrm{LO}}$$

For a thorough review, see: Lee, Prog. Part. Nucl. Phys. 63 (2009) 179

AFQMC + Hybrid Monte Carlo

Substantial investment of supercomputing time ...

CPU time allocations:

- JUQUEEN (FZ Jülich), 30 Mcore-h (project) + > 100 Mcore-h (institutional)

- RWTH cluster (Aachen), 1.3 Mcore-h (project) + "free CPU time" (long queue)



Figure courtesy of Jülich Supercomputer Centre (JSC)

AFQMC results for ¹²C (ground state) Improved NNLO interaction ...





Epelbaum, Krebs, D.L, Meißner, PRL 106 (2011) 192501 Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 109 (2012) 252501

AFQMC – ground and Hoyle states of ${}^{12}C$ Multiple trial wavefunctions ... a = 1.97 fm



Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 109 252501 (2012)

Alpha cluster structure of ^{12}C ...

Ground state







24 rotational orientations

Hoyle state

Production of ¹²**C in red giant stars** Resonant production via ⁸Be and Hoyle state ...

$$r_{3\alpha} = 3^{\frac{3}{2}} N_{\alpha}^3 \left(\frac{2\pi\hbar^2}{M_{\alpha} k_{\rm B} T} \right)^3 \frac{\Gamma_{\gamma}}{\hbar} \, \exp\left(-\frac{\Delta E_{h+b}}{k_{\rm B} T} \right)$$

Is the Universe fine-tuned?



Energy of Hoyle state in ¹²C relative to triple alpha

 $\Delta E_b + \Delta E_h = E_{12}^{\star} - 3E_4$

Experiment: 379.47 ± 0.18 keV

What if the Hoyle state is moved? Calculations of stellar nucleosynthesis ...



Schlattl et al., Astrophys. Space Sci. 291, 27-56 (2004)

 $|\delta(\Delta E_{h+b})| < 100~{\rm keV}$

Anthropic bound on (ad hoc) variation of the Hoyle state

More fundamental description – Chiral EFT Sources of quark mass dependence ...



ChPT:

Figure courtesy of U.-G. Meißner

AFQMC calculation for ⁴He, ⁸Be and ¹²C ...

$$\begin{split} E_{i} &= E_{i}(\tilde{M}_{\pi}, m_{N}(M_{\pi}), \tilde{g}_{\pi N}(M_{\pi}), C_{0}(M_{\pi}), C_{I}(M_{\pi})) \\ \frac{\partial E_{i}}{\partial M_{\pi}}\Big|_{M_{\pi}^{\mathrm{ph}}} &= \left. \frac{\partial E_{i}}{\partial \tilde{M}_{\pi}} \Big|_{M_{\pi}^{\mathrm{ph}}} + x_{1} \left. \frac{\partial E_{i}}{\partial m_{N}} \right|_{m_{N}^{\mathrm{ph}}} + x_{2} \left. \frac{\partial E_{i}}{\partial \tilde{g}_{\pi N}} \right|_{\tilde{g}_{\pi N}^{\mathrm{ph}}} \\ &+ x_{3} \left. \frac{\partial E_{i}}{\partial C_{0}} \right|_{C_{0}^{\mathrm{ph}}} + x_{4} \left. \frac{\partial E_{i}}{\partial C_{I}} \right|_{C_{I}^{\mathrm{ph}}} \\ x_{1} &:= \left. \frac{\partial m_{N}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}} \\ x_{2} &:= \left. \frac{\partial \tilde{g}_{\pi N}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}} = \left. \frac{1}{2F_{\pi}} \left. \frac{\partial g_{A}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}} - \frac{g_{A}}{2F_{\pi}^{2}} \left. \frac{\partial F_{\pi}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}} \\ x_{3} &:= \left. \frac{\partial C_{0}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}}, \quad x_{4} &:= \left. \frac{\partial C_{I}}{\partial M_{\pi}} \right|_{M_{\pi}^{\mathrm{ph}}} \end{split}$$

Parameterization of the short-range terms

Lüscher formula ...

$$\begin{split} p\cot\delta &= \frac{1}{\pi L}S(\eta) \approx -\frac{1}{a}, \qquad \eta := m_N E\left(\frac{L}{2\pi}\right)^2 \\ \bar{A} &= \frac{\partial a^{-1}}{\partial M_{\pi}} = -\frac{1}{\pi L}S'(\eta)\frac{\partial \eta}{\partial M_{\pi}} \\ &-\zeta_s^{-1}\bar{A}_s = \left.\frac{\partial E_s}{\partial \tilde{M}_{\pi}}\right|_{M_{\pi}^{\text{ph}}} + x_1\frac{E_s}{m_N} + x_1\left.\frac{\partial E_s}{\partial m_N}\right|_{m_N^{\text{ph}}} \\ &+ x_2\left.\frac{\partial E_s}{\partial \tilde{g}_{\pi N}}\right|_{\tilde{g}_{\pi N}^{\text{ph}}} + x_3q_s + x_4q_s, \\ &-\zeta_t^{-1}\bar{A}_t = \left.\frac{\partial E_t}{\partial \tilde{M}_{\pi}}\right|_{M_{\pi}^{\text{ph}}} + x_1\frac{E_t}{m_N} + x_1\left.\frac{\partial E_t}{\partial m_N}\right|_{m_N^{\text{ph}}} \\ &+ x_2\left.\frac{\partial E_t}{\partial \tilde{g}_{\pi N}}\right|_{\tilde{g}_{\pi N}^{\text{ph}}} + x_3q_t - 3x_4q_t, \end{split}$$

$$\frac{\partial E_4}{\partial m_{\pi}}\Big|_{m_{\pi}^{\rm phys}} = -0.339(5) \left. \frac{\partial a_s^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} - 0.697(4) \left. \frac{\partial a_t^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} + 0.0380(14)^{+0.008}_{-0.006} + 0.008(14)^{+0.008}_{-0.006} + 0.008(14)^{-0.008}_{-0.008} + 0.008(14)^{-0.008}_{-0.008} + 0.008(14)^{-0.008}_{-0.008} + 0.008(14)^{-0.008}_{-0.008} +$$

⁸Be

$$\frac{\partial E_8}{\partial m_{\pi}}\Big|_{m_{\pi}^{\rm phys}} = -0.794(32) \left. \frac{\partial a_s^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} - 1.584(23) \left. \frac{\partial a_t^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} + 0.089(9)^{+0.017}_{-0.011}$$

¹²C (ground)

$$\frac{\partial E_{12}}{\partial m_{\pi}}\Big|_{m_{\pi}^{\rm phys}} = -1.52(3) \left. \frac{\partial a_s^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} - 2.88(2) \left. \frac{\partial a_t^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} + 0.159(7)^{+0.023}_{-0.018}$$

¹²C (Hoyle)

$$\frac{\partial E_{12}^{\star}}{\partial m_{\pi}}\Big|_{m_{\pi}^{\rm phys}} = -1.588(11) \left. \frac{\partial a_s^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} - 3.025(8) \left. \frac{\partial a_t^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\rm phys}} + 0.178(4)_{-0.021}^{+0.026} + 0.000(4)_{-0.021}^{+0.026}$$

Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 110 (2013) 112502; ibid., arXiv:1303.4856 Berengut et al., Phys. Rev. D 87 (2013) 085018

$$\frac{\partial \Delta E_{h+b}}{\partial m_{\pi}}\Big|_{m_{\pi}^{\text{phys}}} = -0.572(19) \left. \frac{\partial a_s^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\text{phys}}} - 0.933(15) \left. \frac{\partial a_t^{-1}}{\partial m_{\pi}} \right|_{m_{\pi}^{\text{phys}}} + 0.064(6)^{+0.010}_{-0.009}$$

--> Viability of carbon-oxygen based life: $|\delta(\Delta E_{h+b})| < 100~{
m keV}$

$$\left[0.572(19)\bar{A}_s + 0.933(15)\bar{A}_t - 0.064(6) \right] \times \left(\frac{\delta m_q}{m_q} \right) \right| < 0.15\%$$



Epelbaum, Krebs, Lähde, D.L, Meißner, PRL 110 (2013) 112502; ibid., arXiv:1303.4856 Berengut et al., Phys. Rev. D 87 (2013) 085018

Current theoretical knowledge of the quark mass dependence of the S-wave scattering lengths ...



Berengut et al., Phys. Rev. D 87 (2013) 085018

The "end of the world" plot :)



Epelbaum, Krebs, Lähde, Lee, Meißner, Phys. Rev. Lett. 110 (2013) 112502; arXiv:1303.4856

Upcoming results

Spectra of Oxygen-16 and Neon-20

Extension of nuclear lattice EFT up to A = 28

First AFQMC results for non-alpha-cluster nuclei

Extension of chiral NN interaction to N3LO

Effects of lattice spacing and finite volume

Preliminary results for binding energies up to $A = 28 \dots$



(includes contact 4N correction)

Preliminary results for binding energies up to $A = 28 \dots$



Improved NNLO interaction

+ smeared 4N correction